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LHC detector Cosmic-Ray detector Neutrino detector X-ray telescope

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Predicted number of counts = integral over solid angle

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Complicated integrands? → Numerical approaches High-dimensional phase spaces? → Monte Carlo



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$$\mathcal{L} = \bar{\psi}_q^i (i\gamma^\mu) (D_\mu)_{ij} \psi_q^j - m_q \bar{\psi}_q^i \psi_{qi} - \frac{1}{\Delta} F^a_{\mu\nu} F^{a\mu\nu}$$

Quark fields

$$\psi_q^j = \begin{pmatrix} \psi_1 \\ \psi_2 \\ \psi_3 \end{pmatrix}$$

$$\mathcal{L} = \bar{\psi}_q^i (i\gamma^\mu) (D_\mu)_{ij} \psi_q^j - m_q \bar{\psi}_q^i \psi_{qi} - \frac{1}{4} F^a_{\mu\nu} F^{a\mu\nu}$$



$$\mathcal{L} = \bar{\psi}_q^i (i\gamma^\mu) (D_\mu)_{ij} \psi_q^j - m_q \bar{\psi}_q^i \psi_{qi} - \frac{1}{4} F^a_{\mu\nu} F^{a\mu\nu}$$



Gell-Mann Matrices (Ta = ½½)

$$\lambda^{1} = \begin{pmatrix} 0 & 1 & 0 \\ 1 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \ \lambda^{2} = \begin{pmatrix} 0 & -i & 0 \\ i & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \ \lambda^{3} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & -1 & 0 \\ 0 & 0 & 0 \end{pmatrix}, \ \lambda^{4} = \begin{pmatrix} 0 & 0 & 1 \\ 0 & 0 & 0 \\ 1 & 0 & 0 \end{pmatrix}$$
$$\lambda^{5} = \begin{pmatrix} 0 & 0 & -i \\ 0 & 0 & 0 \\ i & 0 & 0 \end{pmatrix}, \ \lambda^{6} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}, \ \lambda^{7} = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & -i \\ 0 & i & 0 \end{pmatrix}, \ \lambda^{8} = \begin{pmatrix} \frac{1}{\sqrt{3}} & 0 & 0 \\ 0 & \frac{1}{\sqrt{3}} & 0 \\ 0 & \frac{1}{\sqrt{3}} & 0 \\ 0 & 0 & -\frac{2}{\sqrt{3}} \end{pmatrix}$$



Color Factors

Color Factors



Color Factors



Color Factors



Color Factors



Color Factors









(from ESHEP lectures by G. Salam)



⁽from ESHEP lectures by G. Salam)

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$$\mathcal{L} = \bar{\psi}_q^i (i\gamma^\mu) (D_\mu)_{ij} \psi_q^j - m_q \bar{\psi}_q^i \psi_{qi} - \frac{1}{4} F^a_{\mu\nu} F^{a\mu\nu}$$

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$$F^a_{\mu\nu} = \partial_\mu A^a_\nu - \partial_\nu A^a_\mu + g_s f^{abc} A^b_\mu A^c_\nu$$

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$$\mathcal{L} = \bar{\psi}_q^i (i\gamma^\mu) (D_\mu)_{ij} \psi_q^j - m_q \bar{\psi}_q^i \psi_{qi} - \frac{1}{4} F^a_{\mu\nu} F^{a\mu\nu}$$

$$F^a_{\mu\nu} = \partial_\mu A^a_\nu - \partial_\nu A^a_\mu + g_s f^{abc} A^b_\mu A^c_\nu$$

 $\begin{array}{c} \mathbf{r} \\ \mathbf{r} \\ \mathbf{r} \\ \mathbf{p} \\ \mathbf{r} \\ \mathbf{p} \\ \mathbf{g} \\ \mathbf$



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& Acum Topological Charge, Data courtesy of M. McGuigan BNL-CSC, T. Izubuchi RIKEN-BNL, and S. Tomov University of Tennessee







Bjorken scaling To first approximation, QCD is SCALE INVARIANT (a.k.a. conformal)

A jet inside a jet inside a jet inside a jet ...

If the strong coupling didn't "run", this would be absolutely true (e.g., N=4 Supersymmetric Yang-Mills)

As it is, α_s only runs slowly (logarithmically) → can still gain insight from fractal analogy



cf. equivalent-photon approximation Weiszäcker, Williams ~ 1934

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Radiation

Radiation

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Radiation

a.k.a. Bremsstrahlung Synchrotron Radiation

Radiation



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$$\propto \frac{1}{2(p_a \cdot p_b)}$$



$$\propto \frac{1}{2(p_a \cdot p_b)} = 00^{a}$$

Gluon j
$$\rightarrow$$
 "soft": Coherence \rightarrow Parton j really emitted by (i,k) "colour antenna"
 $|\mathcal{M}_{F+1}(\dots, i, j, k\dots)|^2 \stackrel{j_g \to 0}{\rightarrow} g_s^2 \mathcal{C} \frac{(p_i \cdot p_k)}{(p_i \cdot p_j)(p_j \cdot p_k)} |\mathcal{M}_F(\dots, i, k, \dots)|^2$

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Can apply this many times \rightarrow nested factorizations

Factorization of Production and Decay:



 $\sim \Lambda_{QCD} \sim 200 \text{ MeV}$

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Factorization 2: PDFs

Hadrons are composite, with time-dependent structure:



Factorization 2: PDFs

Hadrons are composite, with time-dependent structure:

Partons within clouds of further partons, constantly emitted and absorbed



For hadron to remain intact, virtualities $k^2 < M_h^2$ High-virtuality fluctuations suppresed by powers of

$$\frac{\alpha_s M_h^2}{k^2}$$

 M_h : mass of hadron k^2 : virtuality of fluctuation

Factorization 2: PDFs

Hadrons are composite, with time-dependent structure:



\rightarrow Lifetime of fluctuations \sim $1/M_h$

Hard incoming probe interacts over much shorter time scale $\sim 1/Q$

On that timescale, partons ~ frozen

Hard scattering knows nothing of the target hadron apart from the fact that it contained the struck parton

In DIS, there is a formal proof of factorization (Collins, Soper, 1987) Scattering (DIS) (By "deep", we mean $Q^2 >> M_h^2$)

In DIS, there is a formal proof of factorization



$$\sigma^{\ell h} = \sum_{i} \sum_{f} \int dx_i \int d\Phi_f f_{i/h}(x_i, Q_F^2) \, \frac{d\hat{\sigma}^{\ell i \to f}(x_i, \Phi_f, Q_F^2)}{dx_i \, d\Phi_f}$$

In DIS, there is a formal proof of factorization



$$\begin{split} \sigma^{\ell h} = &\sum_{i} \sum_{f} \int dx_{i} \int d\Phi_{f} \, f_{i/h}(x_{i},Q_{F}^{2}) \, \frac{d\hat{\sigma}^{\ell i \rightarrow f}(x_{i},\Phi_{f},Q_{F}^{2})}{dx_{i} \, d\Phi_{f}} \\ & \text{Sum over} \\ & \text{Initial (i)} \\ & \text{and final (f)} \\ & \text{parton flavors} \end{split}$$

In DIS, there is a formal proof of factorization



$$\begin{split} \sigma^{\ell h} = &\sum_{i} \sum_{f} \int dx_{i} \int d\Phi_{f} f_{i/h}(x_{i}, Q_{F}^{2}) \frac{d\hat{\sigma}^{\ell i \to f}(x_{i}, \Phi_{f}, Q_{F}^{2})}{dx_{i} d\Phi_{f}} \\ & \text{Sum over} \qquad \Phi_{f} \\ & \text{Initial (i)} \qquad = \text{Final-state} \\ & \text{and final (f)} \qquad \text{phase space} \end{split}$$

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In DIS, there is a formal proof of factorization



$$\sigma^{\ell h} = \sum_{i} \sum_{f} \int dx_{i} \int d\Phi_{f} f_{i/h}(x_{i}, Q_{F}^{2}) \frac{d\hat{\sigma}^{\ell i \to f}(x_{i}, \Phi_{f}, Q_{F}^{2})}{dx_{i} d\Phi_{f}}$$
Sum over Φ_{f} $f_{i/h}$
Initial (i) = Final-state = PDFs and final (f) phase space Assumption: $Q^{2} = Q_{F}^{2}$
Differential partonic Hard-scattering Matrix Element(s)

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Initial (i) = Final-state = PDFs and final (f) phase space Assumption: $Q^{2} = Q_{F}^{2}$
Differential partonic Hard-scattering Matrix Element(s)

It's just another crossing



Factorization

Why is Fixed Order QCD not enough?

: It requires all resolved scales >> Λ_{QCD} **AND** no large hierarchies

Trivially untrue for QCD

We're colliding, and observing, hadrons \rightarrow small scales We want to consider high-scale processes \rightarrow large scale differences

→ A Priori, no perturbatively calculable observables in hadron-hadron collisions

Factorization

Why is Fixed Order QCD not enough?

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Trivially untrue for QCD

We're colliding, and observing, hadrons \rightarrow small scales We want to consider high-scale processes \rightarrow large scale differences

$$\frac{\mathrm{d}\sigma}{\mathrm{d}X} = \sum_{a,b} \sum_{f} \int_{\hat{X}_{f}} f_{a}(x_{a}, Q_{i}^{2}) f_{b}(x_{b}, Q_{i}^{2}) \frac{\mathrm{d}\hat{\sigma}_{ab \to f}(x_{a}, x_{b}, f, Q_{i}^{2}, Q_{f}^{2})}{\mathrm{d}\hat{X}_{f}} D(\hat{X}_{f} \to X, Q_{i}^{2}, Q_{f}^{2})$$

PDFs: needed to compute inclusive cross sections

FFs: needed to compute (semi-)exclusive cross sections

Factorization

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PDFs: needed to compute inclusive cross sections

FFs: needed to compute (semi-)exclusive cross sections

Resummed pQCD: All resolved scales >> Λ_{QCD} **AND** X Infrared Safe

^{*)}pQCD = perturbative QCD

P. Skands Event Generator Physics Lecture 2 Terascale Monte Carlo School DESY Hamburg March 2014

Introduction to Monte Carlo

Recall: Scattering Experiments



LHC detector Cosmic-Ray detector Neutrino detector X-ray telescope

→ Integrate differential cross sections over specific phase-space regions

Predicted number of counts = integral over solid angle

 $N_{\rm count}(\Delta\Omega) \propto \int_{\Delta\Omega} \mathrm{d}\Omega \frac{\mathrm{d}\sigma}{\mathrm{d}\Omega}$

In particle physics:

Integrate over all quantum histories (+ interferences) Complicated integrands? → Numerical approaches High-dimensional phase spaces? → Monte Carlo

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Integrate over all quantum histories (+ interferences) Complicated integrands? → Numerical approaches High-dimensional phase spaces? → Monte Carlo

Part of $Z \rightarrow 4$ jets ...

5.3 Four-parton tree-level antenna functions

The tree-level four-parton quark-antiquark antenna contains three final states: quarkgluon-gluon-antiquark at leading and subleading colour, A_4^0 and \tilde{A}_4^0 and quark-antiquarkquark-antiquark for non-identical quark flavours B_4^0 as well as the identical-flavour-only contribution C_4^0 . The quark-antiquark-quark-antiquark final state with identical quark flavours is thus described by the sum of antennae for non-identical flavour and identicalflavour-only. The antennae for the $qgg\bar{q}$ final state are:

$$\begin{split} & A_{4}^{0}(1_{q},3_{g},4_{g},2_{\tilde{q}}) = a_{4}^{0}(1,3,4,2) + a_{4}^{0}(2,4,3,1), \qquad (5.27) \\ & \tilde{A}_{4}^{0}(1_{q},3_{g},4_{g},2_{\tilde{q}}) = \tilde{a}_{4}^{0}(1,3,4,2) + \tilde{a}_{4}^{0}(2,4,3,1) + \tilde{a}_{4}^{0}(1,4,3,2) + \tilde{a}_{4}^{0}(2,3,4,1), \quad (5.28) \\ & a_{4}^{0}(1,3,4,2) = \frac{1}{s_{1234}} \Biggl\{ \frac{1}{2s_{13}s_{24}s_{134}} \Bigl[2s_{12}s_{14} + 2s_{12}s_{23} + 2s_{12}^{2} + s_{14}^{2} + s_{23}^{2} \Bigr] \\ & + \frac{1}{2s_{13}s_{24}s_{134}s_{234}} \Bigl[3s_{12}s_{23}^{2} - 4s_{12}^{2}s_{34} + 2s_{12}^{3} - s_{34}^{3} \Bigr] \\ & + \frac{1}{s_{13}s_{24}s_{134}} \Bigl[3s_{12}s_{23} - 3s_{12}s_{34} + 4s_{12}^{2} - s_{23}s_{34} + s_{23}^{2} + s_{34}^{2} \Bigr] \\ & + \frac{3}{2s_{13}s_{24}} \Bigl[2s_{12} + s_{14} + s_{23} \Bigr] + \frac{1}{s_{13}s_{34}} \Bigl[4s_{12} + 3s_{23} + 2s_{24} \Bigr] \\ & + \frac{3}{2s_{13}s_{24}} \Bigl[2s_{12} + s_{14} + s_{23} \Bigr] + \frac{1}{s_{13}s_{134}} \Bigl[3s_{12}s_{24} + 6s_{12}s_{34} - 4s_{12}^{2} - 3s_{24}s_{34} - s_{24}^{2} - 3s_{34}^{2} \Bigr] \\ & + \frac{1}{s_{13}s_{13}s_{234}} \Bigl[2s_{12} + s_{14} + s_{23} \Bigr] + \frac{1}{s_{13}s_{134}} \Bigl[4s_{12} + 3s_{23} + 2s_{24} \Bigr] \\ & + \frac{1}{s_{13}s_{13}s_{134}} \Bigl[-6s_{12} - 3s_{23} - s_{24} + 2s_{34} \Bigr] \\ & + \frac{1}{s_{13}s_{134}s_{234}} \Bigl[2s_{12}s_{14} + 2s_{12}s_{23} + 2s_{12}^{2} - 3s_{24}s_{34} - s_{24}^{2} - 3s_{34}^{2} \Bigr] \\ & + \frac{1}{s_{24}s_{43}s_{4}s_{134}} \Bigl[2s_{12}s_{14} + 2s_{12}s_{23} + 2s_{12}^{2} + 2s_{14}s_{23} + s_{14}^{2} + s_{23}^{2} \Bigr] \\ & + \frac{1}{s_{24}s_{13}s_{134}} \Bigl[-4s_{12} - s_{14} - s_{23} + s_{34} \Bigr] + \frac{1}{s_{34}^{2}} \Bigl[s_{12} + 2s_{13} - 2s_{14} - s_{34} \Bigr] \\ & + \frac{1}{s_{24}s_{134}} \Bigl[2s_{12}s_{14}^{2} + 2s_{14}^{2}s_{23} + 2s_{14}^{2}s_{24} \Bigr] - \frac{2s_{12}s_{14}s_{24}}{s_{23}^{2}s_{134}s_{234}} \\ & + \frac{1}{s_{34}^{2}s_{134}} \Bigl[-2s_{12}s_{14} - 4s_{14}s_{24} + 2s_{14}^{2} \Biggr] \\ & + \frac{1}{s_{34}^{2}s_{134}s_{234}} \Bigl[-2s_{12}s_{14} - 4s_{12}^{2} + 2s_{14}s_{24} - s_{14}^{2} - s_{24}^{2} \Biggr] \\ & + \frac{1}{s_{34}^{2}s_{134}s_{234}} \Bigl[-2s_{12}s_{14} - 4s_{12}^{2} + 2s_{14}s_{24} - s_{14}^{2} - s_{24}^{2} \Biggr] \\ & + \frac{1}{s_{34}^{2}s_{134}s_{234}} \Bigl[-2s_{12}s_{14} - 4s_{12}^{2} + 2s_{14}s_{24} - s_{14}^{2}$$

$$\begin{split} \bar{a}_{4}^{0}(1,3,4,2) &= \frac{1}{s_{1234}} \Biggl\{ \frac{1}{s_{13}s_{24}s_{134}s_{234}} \Biggl[\frac{3}{2}s_{12}s_{34}^{2} - 2s_{12}^{2}s_{34} + s_{12}^{3} - \frac{1}{2}s_{34}^{3} \Biggr] \\ &+ \frac{1}{s_{13}s_{24}s_{134}} \Biggl[3s_{12}s_{23} - 3s_{12}s_{34} + 4s_{12}^{2} - s_{23}s_{34} + s_{23}^{2} + s_{34}^{2} \Biggr] \\ &+ \frac{s_{13}^{3}}{s_{13}s_{24}(s_{13} + s_{23})(s_{14} + s_{24})} + \frac{1}{s_{13}s_{24}(s_{13} + s_{23})} \Biggl[\frac{1}{2}s_{12}s_{14} + s_{12}^{2} \Biggr] \\ &+ \frac{1}{s_{13}s_{24}(s_{14} + s_{24})} \Biggl[\frac{1}{2}s_{12}s_{23} + s_{12}^{2} \Biggr] + \frac{1}{s_{13}s_{24}} \Biggl[3s_{12} + \frac{3}{2}s_{14} + \frac{3}{2}s_{23} \Biggr] \\ &+ \frac{1}{s_{13}s_{134}^{2}} \Biggl[s_{12}s_{34} + s_{23}s_{34} + s_{24}s_{34} \Biggr] + \frac{2s_{12}^{3}}{s_{13}s_{134}s_{234}(s_{13} + s_{23})} \Biggr] \\ &+ \frac{1}{s_{13}s_{134}^{2}} \Biggl[s_{12}s_{34} - s_{24}s_{34} - 2s_{34}^{2} \Biggr] \\ &+ \frac{1}{s_{13}s_{134}(s_{13} + s_{23})} \Biggl[s_{12}s_{24} + s_{12}s_{34} - 2s_{12}^{2} \Biggr] \\ &+ \frac{1}{s_{13}s_{134}(s_{13} + s_{23})} \Biggl[s_{12}s_{24} + s_{12}s_{34} + 2s_{12}^{2} \Biggr] \\ &+ \frac{1}{s_{13}s_{134}} \Biggl[-2s_{12} - 2s_{14} + s_{24} + 2s_{34} \Biggr] \\ &+ \frac{2s_{12}^{3}}{s_{13}(s_{13} + s_{23})(s_{14} + s_{24})(s_{13} + s_{14})} \Biggr] \\ &+ \frac{1}{s_{13}(s_{13} + s_{23})(s_{14} + s_{24})(s_{13} + s_{14})} \Biggr] \Biggr] \\ &+ \frac{1}{s_{13}(s_{13} + s_{23})(s_{13} + s_{14})} \Biggr[s_{12}s_{24} + 2s_{12}^{2} \Biggr] \\ &+ \frac{1}{s_{13}(s_{13} + s_{23})(s_{13} + s_{14})} \Biggr[s_{12}s_{24} + 2s_{12}^{2} \Biggr] \\ &+ \frac{1}{s_{13}(s_{13} + s_{23})(s_{13} + s_{14})} \Biggr[s_{12}s_{23} + 2s_{12}^{2} \Biggr] \\ &+ \frac{1}{s_{13}(s_{13} + s_{24})(s_{13} + s_{14})} \Biggr[s_{12}s_{23} + 2s_{12}^{2} \Biggr] \\ &+ \frac{1}{s_{13}(s_{13} + s_{24})(s_{13} + s_{14})} \Biggr[s_{12}s_{23} + 2s_{12}^{2} \Biggr] \\ &+ \frac{1}{s_{13}(s_{13} + s_{24})(s_{13} + s_{14})} \Biggr[s_{12}s_{23} + 2s_{12}^{2} \Biggr] \\ &+ \frac{1}{s_{13}(s_{13} + s_{14})} - \frac{2}{s_{13}}} \Biggr] \Biggr]$$

+ Additional Subleading Terms ...

(5.29)

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Part of $Z \rightarrow 4$ jets ...

5.3 Four-parton tree-level antenna functions

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$$\begin{split} \tilde{a}_{4}^{0}(1,3,4,2) &= \frac{1}{s_{1234}} \left\{ \frac{1}{s_{13}s_{24}s_{134}s_{234}} \left[\frac{3}{2}s_{12}s_{34}^{2} - 2s_{12}^{2}s_{34} + s_{12}^{2} - \frac{1}{2}s_{34}^{3} \right] \\ &+ \frac{1}{s_{13}s_{24}s_{134}} \left[3s_{12}s_{23} - 3s_{12}s_{34} + 4s_{12}^{2} - s_{23}s_{34} + s_{23}^{2} + s_{34}^{2} \right] \\ &+ \frac{1}{s_{13}s_{24}s_{134}} \left[3s_{12}s_{23} - 3s_{12}s_{34} + 4s_{12}^{2} - s_{23}s_{34} + s_{23}^{2} + s_{34}^{2} \right] \\ &+ \frac{1}{s_{13}s_{24}s_{134}} \left[1 + \frac{1}{s_{13}s_{134}} + \frac{1}{s_{13}s_{134}} \right] \left[s_{12}s_{23} - s_{24} + s_{12}s_{34} + 2s_{12}^{2} \right] \\ &+ \frac{1}{s_{13}s_{134}} \left[-s_{23} - s_{24} + 2s_{34} \right] + \frac{1}{s_{13}s_{234}} \left[s_{12}s_{14} + s_{12}s_{34} + 2s_{12}^{2} \right] \\ &+ \frac{1}{s_{13}s_{134}} \left[-2s_{12} - 2s_{14} + s_{24} + 2s_{34} \right] \\ &+ \frac{2s_{12}^{2}}{s_{13}(s_{13} + s_{23})(s_{14} + s_{24})(s_{13} + s_{14})} \\ &+ \frac{1}{s_{13}(s_{13} + s_{23})(s_{14} + s_{24})(s_{13} + s_{14})} \\ &+ \frac{1}{s_{13}(s_{13} + s_{23})(s_{14} + s_{14})} \left[s_{12}s_{23} + 2s_{12}^{2} \right] \\ &+ \frac{1}{s_{13}(s_{13} + s_{24})(s_{13} + s_{14})} \\ &+ \frac{1}{s_{13}(s_{13} + s_{24})(s_{13} + s_{14})} \left[s_{12}s_{23} + 2s_{12}^{2} \right] \\ &+ \frac{1}{s_{13}(s_{13} + s_{24})(s_{13} + s_{14})} \left[s_{12}s_{23} + 2s_{12}^{2} \right] \\ &+ \frac{1}{s_{13}(s_{13} + s_{24})(s_{13} + s_{14})} \left[s_{12}s_{23} + 2s_{12}^{2} \right] \\ &+ \frac{1}{s_{13}(s_{13} + s_{24})(s_{13} + s_{14})} \left[s_{12}s_{23} + 2s_{12}^{2} \right] \\ &+ \frac{1}{s_{13}(s_{13} + s_{24})(s_{13} + s_{14})} \left[s_{12}s_{23} + 2s_{12}^{2} \right] \\ &+ \frac{1}{s_{13}(s_{13} + s_{24})(s_{13} + s_{14})} \left[s_{12}s_{23} + 2s_{12}^{2} \right] \\ &+ \frac{1}{s_{13}(s_{13} + s_{24})} \left[s_{12} - s_{34} \right] \\ &+ \frac{1}{s_{13}(s_{13} + s_{14})} - \frac{2}{s_{13}} + \frac{1}{s_{134}^{2}} \left[s_{12} + s_{23} + s_{24} \right] \\ &+ \frac{1}{s_{13}(s_{13} + s_{14})} - \frac{2}{s_{13}} + \frac{1}{s_{134}^{2}} \left[s_{12} - s_{23} + s_{24} \right] \\ &+ \frac{1}{s_{13}(s_{12} + s_{24}} + \frac{1}{s_{12}} \left[s_{12} - s_{23} + s_{24} \right] \\ &+ \frac{1}{s_{13}(s_{12} + s_{24})} \left[s_{12} - s_{23} + s_{24} \right] \\ &+ \frac{1}{s_{13}(s_{13} + s_{14})} - \frac{2}{s_{13}} + \frac{1}{s_{134}^{2}} \left[s_{12} - s_{23} + s_{24} \right] \\$$

+ Additional Subleading Terms ...

(5.29)

Part of $Z \rightarrow 4$ jets ...

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Now compute and add the quantum corrections ...

$$+\frac{1}{s_{13}(s_{13}+s_{23})(s_{14}+s_{24})(s_{13}+s_{14})} + \frac{1}{s_{13}(s_{13}+s_{23})(s_{13}+s_{14})} \left[s_{12}s_{24}+2s_{12}^{2}\right] + \frac{1}{s_{13}(s_{14}+s_{24})(s_{13}+s_{14})} \left[s_{12}s_{23}+2s_{12}^{2}\right] + \frac{2s_{12}}{s_{13}(s_{13}+s_{14})} - \frac{2}{s_{13}} + \frac{1}{s_{134}^{2}} \left[s_{12}+s_{23}+s_{24}\right] + \frac{1}{s_{134}s_{234}} \left[s_{12}-s_{34}\right] + \frac{1}{s_{134}} + \mathcal{O}(\epsilon) \right\}.$$
(5.30)

+ Additional Subleading Terms ...

(5.29)

 $-2s_{12}^2$

Part of $Z \rightarrow 4$ jets ...

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 $\tilde{a}_{4}^{0}(1,3,4,2) = \frac{1}{s_{1234}} \left\{ \frac{1}{s_{13}s_{24}s_{134}s_{234}} \left[\frac{3}{2}s_{12}s_{34}^2 - 2s_{12}^2s_{34} + s_{12}^3 - \frac{1}{2}s_{34}^3 \right] \right\}$ $+\frac{1}{s_{13}s_{24}s_{134}} \begin{bmatrix} 3s_{12}s_{23} - 3s_{12}s_{34} + 4s_{12}^2 - s_{23}s_{34} + s_{23}^2 + s_{34}^2 \end{bmatrix}$ This is one of the simplest processes ... computed at lowest order in the theory. *s*₁₃*s*₁₃₄*s*₂₃₄ ^L $+\frac{1}{s_{13}s_{134}(s_{13}+s_{23})}\left[s_{12}s_{24}+s_{12}s_{34}+2s_{12}^2\right]$ $-2s_{12}^2$ Now compute and add the quantum corrections ... $+\frac{12}{s_{13}(s_{13}+s_{23})(s_{14}+s_{24})(s_{13}+s_{14})}$ $+\frac{1}{s_{13}(s_{13}+s_{23})(s_{13}+s_{14})}\left[s_{12}s_{24}+2s_{12}^2\right]$ Then maybe worry about simulating the detector

too ...

(5.30)

+ Additional Subleading Terms ...

(5.29)

Riemann Sums



Numerical Integration in 1D

Midpoint (rectangular) Rule:

Fixed-Grid n-point Quadrature Rules

Divide into N "bins" of size Δ Approximate f(x) \approx constant in each bin Sum over all rectangles inside your region



Numerical Integration in 1D

Trapezoidal Rule:

Fixed-Grid n-point Quadrature Rules

Approximate $f(x) \approx \text{linear}$ in each bin Sum over all trapeziums inside your region



Numerical Integration in 1D

Simpson's Rule:

Fixed-Grid n-point Quadrature Rules

Approximate $f(x) \approx quadratic$ in each bin Sum over all "Simpsons" inside your region



... and so on for higher n-point rules ...

Convergence Rate

The most important question:

Recipes

How long do I have to wait? How many evaluations do I need to calculate for a given precision?

Uncertainty (after n evaluations)	n _{eval} / bin	Approx Conv. Rate (in 1D)
Trapezoidal Rule (2-point)	2	I/N ²
Simpson's Rule (3-point)	3	I/N ⁴
m-point (Gauss quadrature)	m	I/N ^{2m-I}
See, e.g., Numerical	See, e.g., F. James, "Monte	

Carlo Theory and Practice"

Higher Dimensions

Fixed-Grid (Product) Rules scale exponentially with D

m-point rule in 1 dimension



→ m function evaluations per bin

... in 2 dimensions



 \rightarrow m² evaluations per bin

Higher Dimensions

Fixed-Grid (Product) Rules scale exponentially with D

m-point rule in 1 dimension



 \rightarrow m function evaluations per bin

... in 2 dimensions



E.g., to evaluate a 12-point rule in 10 dimensions, need 1000 billion evaluations per bin

Convergence Rate

+ Convergence is slower in higher Dimensions!

→ More points for less precision

Recipes

Uncertainty (after n evaluations)	n _{eval} / bin	Approx Conv. Rate (in D dim)
Trapezoidal Rule (2-point)	2 ^D	l/n ^{2/D}
Simpson's Rule (3-point)	3 D	l/n ^{4/D}
m-point (Gauss rule)	m ^D	I/n ^{(2m-I)/D}
See, e.g., Numerical	See, e.g., F. Ja	mes, "Monte

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A Monte Carlo technique: is any technique making use of random numbers to solve a problem



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Convergence:

<u>Calculus:</u> {A} converges to B if an n exists for which $|A_{i>n} - B| < \varepsilon$, for any $\varepsilon > 0$

Monte Carlo: {A} converges to B if n exists for which the probability for |A_{i>n} - B| < ε, for any ε > 0, is > P, for any P[0<P<1]</p>



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Monte Carlo: {A} converges to B if n exists for which the probability for |A_{i>n} - B| < ε, for any ε > 0, is > P, for any P[0<P<1] "This risk, that convergence is only given with a certain probability, is inherent in Monte Carlo calculations and is the reason why this technique was named after the world's most famous gambling casino. Indeed, the name is doubly appropriate because the style of gambling in the Monte Carlo casino, not to be confused with the noisy and tasteless gambling houses of Las Vegas and Reno, is serious and sophisticated."

F. James, "Monte Carlo theory and practice", Rept. Prog. Phys. 43 (1980) 1145

Convergence

MC convergence is Stochastic!

in <u>any</u> dimension

ted dot size red dot size red dot spacing

Uncertainty (after n function evaluations)	n _{eval} / bin	Approx Conv. Rate (in ID)	Approx Conv. Rate (in D dim)
Trapezoidal Rule (2-point)	2 D	l/n²	I/n ^{2/D}
Simpson's Rule (3-point)	3 D	I/n ⁴	I/n ^{4/D}
m-point (Gauss rule)	m ^D	l/n ^{2m-l}	l/n ^{(2m-1)/D}
Monte Carlo	I	I/n ^{1/2}	l/n ^{1/2}

+ many ways to optimize: stratification, adaptation, …
+ gives "events" → iterative solutions,
+ interfaces to detector simulation & propagation codes

Random Numbers

(apologies, I did not have much time to adapt these slides)

You want: to know the area of this shape:


(apologies, I did not have much time to adapt these slides)

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Assume you know the area of <u>this</u> shape: πR^2 (an overestimate)

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You want: to know the area of this shape:

Now get a few friends, some balls, and throw random shots inside the circle (but be careful to make your shots truly random)



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Count how many shots hit the shape inside and how many miss



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Count how many shots hit the shape inside and how many miss



Assume you know the area of this shape: ΠR^2 (an overestimate)

(apologies, I did not have much time to adapt these slides)

You want: to know the area of this shape:

Now get a few friends, some balls, and throw random shots inside the circle (but be careful to make your shots truly random)

Count how many shots hit the shape inside and how many miss



Assume you know the area of <u>this</u> shape: πR² (an overestimate)



G. Leclerc, Comte de Buffon (1707-1788)

 $A_{a} \approx N_{hit}/N_{miss} \times \Pi R^{2}$

I will not tell you how to *write* a Random-number generator (interesting topic & history in its own right)

Instead, I <u>assume</u> that you can write a computer code and link to a random-number generator, from a library

E.g., ROOT includes one that you can use if you like.

PYTHIA also includes one

From the PYTHIA 8 HTML documentation, under <u>"Random Numbers"</u>:

Random numbers R uniformly distributed in 0 < R < 1 are obtained with

Pythia8::Rndm::flat();

+ Other methods for exp, x^*exp , 1D Gauss, 2D Gauss.

Example: Number of Terascale school students who will get hit by a car this week

Complicated Function:

Time-dependent

Traffic density during day, week-days vs week-ends

(i.e., non-trivial time evolution of system)

No two students are the same

Need to compute probability for each and sum

(simulates having several distinct types of "evolvers")

Multiple outcomes:

Hit → keep walking, or go to hospital? Multiple hits = Product of single hits, or more complicated?

Monte Carlo Approach

Approximate Traffic

- Simple overestimate:
 - highest recorded density of most careless drivers, driving at highest recorded speed



Approximate Student

. . .

by most completely reckless and accident-prone student (wandering the streets lost in thought after these lectures ...)

This extreme guess will be the equivalent of our simple overestimate from before:

Off we go...

Throw random accidents according to:

 $n_{\rm stud}$ $\sum \alpha_i(x,t) \rho_i(x,t) \rho_c(x,t)$ Student-Car Density of Density of Student i hit rate Cars Sum over students

Off we go...

Throw random accidents according to:



Too

Difficult

Off we go...

Throw random accidents according to:



P. Skands

Off we go...

Throw random accidents according to:



Simple

Overestimate

P. Skands

Off we go...

Throw random accidents according to:



 $\mathsf{R} = (t_e - t_0) \Delta x$

$\alpha_{\max} n_{\text{stud}} \rho_{c\max}$

Hit rate for most accident-prone student Rush-hour density of cars



Off we go...

Throw random accidents according to:



of cars

(Also generate trial *x_e*, e.g., uniformly in circle around DESY) (Also generate trial *i*; a random student gets hit)

student

Accept trial hit (i,x,t) with probability

Prob(accept) =
$$\frac{\alpha_i(x,t) \ \rho_i(x,t) \ \rho_c(x,t)}{\alpha_{\max} \ n_{\text{stud}} \ \rho_{c\max}}$$

Using the following:

 ρ_c : actual density of cars at location x at time t ρ_i : actual density of student i at location x at time t α_i : The actual "hit rate" (OK, not really known, but can make one up)

Accept trial hit (i,x,t) with probability

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Using the following:

 ho_c : actual density of cars at location x at time t ho_i : actual density of student i at location x at time t ho_i : The actual "hit rate" (OK, not really known, but can make one up)

→ True number = number of accepted hits (note: we didn't really treat multiple hits ... → Markov Chain)

Importance Sampling



Peaked Functions



Peaked Functions



Peaked Functions



Precision on integral dominated by the points with $f \approx f_{max}$ (*i.e.*, peak regions)

→ slow convergence if high, narrow peaks





→ Make it twice as likely to throw points in the peak

Choose:

	[0,1] →	Region A
For:	[1,2] →	Region B
6*R1	∈ [2,4] →	Region C
-=-	[4,5] →	Region D
	[5,6] →	Region E

→ faster convergence
 for same number
 of function evaluations

Adaptive Sampling



→ Can even design
 algorithms to
 do this automatically
 as they run
 (not covered here)

→ Adaptive sampling

Importance Sampling



→ or throw points according to some smooth peaked function for which you have, or can construct, a random number generator (here: Gauss)

E.g., VEGAS algorithm, by G. Lepage

1) You are inputting knowledge: obviously need to know where the peaks are to begin with ... (say you know, e.g., the location and width of a resonance)

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2) Stratified sampling increases efficiency by combining fixed-grid methods with the MC method, with further gains from adaptation

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2) Stratified sampling increases efficiency by combining fixed-grid methods with the MC method, with further gains from adaptation

3) Importance sampling:

$$\int_{a}^{b} f(x) dx = \int_{a}^{b} \frac{f(x)}{g(x)} dG(x)$$

Effectively does flat MC with changed integration variables Fast convergence if $f(x)/g(x) \approx 1$

The Veto Algorithm



Miss

Take your system

Set of radioactive nuclei Set of hard scattering processes Set of resonances that are going to decay Set of particles coming into your detector Set of cosmic photons traveling across the galaxy Set of molecules



...

Take your system

Generate a "trial" (event/decay/interaction/...)

- Not easy to generate random numbers distributed according to exactly the right distribution?
- May have complicated dynamics, interactions ...
- → use a simpler "trial" distribution

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Flat with some stratification

Or importance sample with simple overestimating function (for which you can generate random #s)

Take your system

Generate a "trial" (event/decay/interaction/...)
Accept trial with probability f(x)/g(x)
 f(x) contains all the complicated dynamics
 g(x) is the simple trial function
If accept: replace with new system state
If reject: keep previous system state

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And keep going: generate next trial ...



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no dependence on g in final result - only affects convergence rate

And keep going: generate next trial ...


How we do Monte Carlo

Take your system Generate a "trial" (event/dec Accept trial with probability f(x f(x) contains all the complicated g(x) is the simple trial function If accept: replace with new sys If reject: keep previous system

no dependence on g ir result - only affect convergence rate

Sounds deceptively simple, but ...

with it, you can integrate

arbitrarily complicated functions (in particular chains of nested functions), over arbitrarily complicated regions, in arbitrarily many dimensions ...

And keep going: generate next trial ...



Summary - Lecture 2

Quantum Scattering Problems are common to many areas of physics: To compute expectation value of observable: integrate over phase space

Complicated functions → Numerical Integration

High Dimensions → Monte Carlo (stochastic) convergence is fastest + Additional power by stratification and/or importance sampling



Recommended Reading

F. James Monte Carlo Theory and Practice Rept.Prog.Phys.43 (1980) p.1145

S.Weinzierl Topical lectures given at the Research School Subatomic physics, Amsterdam, June 2000 Introduction to Monte Carlo Methods e-Print: hep-ph/0006269

P. Skands Introduction to QCD (TASI 2012) arXiv:1207.2389

Naively, QCD radiation suppressed by $\alpha_s \approx 0.1$

Truncate at fixed order = LO, NLO, ...

But beware the jet-within-a-jet-within-a-jet ...

Example:

SUSY pair production at 14 TeV, with $M_{\text{SUSY}}\approx 600~\text{GeV}$

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LHC - sps1a - m~600 GeV

FIXED ORDER pQCD	$\sigma_{\rm tot}[{\rm pb}]$	$ ilde{g} ilde{g}$	$\tilde{u}_L \tilde{g}$	$\tilde{u}_L \tilde{u}_L^*$	$\tilde{u}_L \tilde{u}_L$	TT
$p_{T,j} > 100 \text{ GeV}$	σ_{0j}	4.83	5.65	0.286	0.502	1.30
inclusive X + 1 "jet" —	$\rightarrow \sigma_{1j}$	2.89	2.74	0.136	0.145	0.73
inclusive X + 2 "jets" ⁻	$\rightarrow \sigma_{2j}$	1.09	0.85	0.049	0.039	0.26

Plehn, Rainwater, PS PLB645(2007)217

σ for X + jets much larger than naive estimate

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$p_{T,j} > 100 { m GeV}$ inclusive X + 1 "jet" — inclusive X + 2 "jets" —	σ_{0j} $\rightarrow \sigma_{1j}$ $\rightarrow \sigma_{2j}$	4.83 2.89 1.09	5.65 2.74 0.85	$0.286 \\ 0.136 \\ 0.049$	$0.502 \\ 0.145 \\ 0.039$	$1.30 \\ 0.73 \\ 0.26$	σ for X + jets much larger than naive estimate
$p_{T,j} > 50 \text{ GeV}$	$\sigma_{0j} \ \sigma_{1j} \ \sigma_{2j}$	4.83 5.90 4.17	5.65 5.37 3.18	0.286 0.283 0.179	0.502 0.285 0.117	1.30 1.50 1.21	σ for 50 GeV jets ≈ larger than total cross section → not under control
L			(Co	omputed with	h SUSY-Mao	dGraph)	

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Example: 100 GeV can be "soft" at the LHC

→ More on this in lectures on Jets, Monte Carlo, and Matching

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P. Skands

Scaling Violation

Real QCD isn't conformal

The coupling runs logarithmically with the energy scale



Scaling Violation

Real QCD isn't conformal

The coupling runs logarithmically with the energy scale



Asymptotic freedom in the ultraviolet

Scaling Violation

Real QCD isn't conformal

The coupling runs logarithmically with the energy scale

$$\begin{split} Q^2 \frac{\partial \alpha_s}{\partial Q^2} &= \beta(\alpha_s) \qquad \beta(\alpha_s) = -\alpha_s^2 (b_0 + b_1 \alpha_s + b_2 \alpha_s^2 + \dots) \ , \\ b_0 &= \frac{11C_A - 2n_f}{12\pi} \quad b_1 = \frac{17C_A^2 - 5C_A n_f - 3C_F n_f}{24\pi^2} = \frac{153 - 19n_f}{24\pi^2} \\ \text{I-Loop } \beta \text{ function coefficient} \qquad \text{2-Loop } \beta \text{ function coefficient} \qquad b_2 = b_3 = known \end{split}$$

Asymptotic freedom in the ultraviolet

Confinement (IR slavery?) in the infrared

"What this year's Laureates discovered was something that, at first sight, seemed completely contradictory. The interpretation of their mathematical result was that the closer the quarks are to each other, the *weaker* is the 'colour charge'. When the quarks are really close to each other, the force is so weak that they behave almost as free particles. This phenomenon is called 'asymptotic freedom'. The converse is true when the quarks move apart: the force becomes stronger when the distance increases."



Nobelprize.org

The Official Web Site of the Nobel Prize

The Nobel Prize in Physics 2004 David J. Gross, H. David Politzer, Frank Wilczek



David J. GrossH. David PolitzerFrank WilczekThe Nobel Prize in Physics 2004 was awarded jointly to David J. Gross, H. David Politzer and FrankWilczek "for the discovery of asymptotic freedom in the theory of the strong interaction".

Photos: Copyright © The Nobel Foundation

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- *1 each other, the force is so weak that they behave almost as free particles. This phenomenon is called 'asymptotic freedom'. The converse is true when the quarks move apart:
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Photos: Copyright © The Nobel Foundation



^{*1} The force still goes to ∞ as $r \rightarrow 0$ (Coulomb potential), just less slowly

^{*2} The potential grows linearly as $r \rightarrow \infty$, so the force actually becomes constant (even this is only true in "quenched" QCD. In real QCD, the force eventually vanishes for r>>1 fm)

QED:

Vacuum polarization

→ Charge screening



Quark Loops

→ Also charge screening





But only dominant if > 16 flavors!

QED:

Vacuum polarization

→ Charge screening

QCD:



Gluon Loops Dominate if \leq 16 flavors





Spin-I → Opposite Sign

UV and IR



At low scales

Coupling $\alpha_s(Q)$ actually runs rather fast with Q

Perturbative solution diverges at a scale Λ_{QCD} somewhere below

 \approx I GeV

So, to specify the strength of the strong force, we usually give the value of α_s at a unique reference scale that everyone agrees on: M_Z

The Fundamental Parameter(s)



... + nf and quark masses

61

The Fundamental Parameter(s)



... + nf and quark masses

The Fundamental Parameter(s)



... And all its cousins

 $\Lambda^{(3)} \Lambda^{(4)} \Lambda^{(5)} \Lambda_{CMW} \Lambda_{FSR} \Lambda_{ISR} \Lambda_{MPI}, \dots$

... + nf and quark masses

Uncalculated Orders

Naively $O(\alpha_s)$ - True in e⁺e⁻!



Uncalculated Orders

Naively O(α_s) - True in e⁺e⁻!



Generally larger in hadron collisions

- Typical "K" factor in pp (= σ_{NLO}/σ_{LO}) $\approx 1.5 \pm 0.5$
- Why is this? Many pseudoscientific explanations

Uncalculated Orders

Naively O(α_s) - True in e⁺e⁻!

$$\sigma_{\rm NLO}(e^+e^- \to q\bar{q}) = \sigma_{\rm LO}(e^+e^- \to q\bar{q}) \left(1 + \left(\frac{\alpha_s(E_{\rm CM})}{\pi} + \mathcal{O}(\alpha_s^2)\right)\right)$$

Generally larger in hadron collisions

Typical "K" factor in pp (= σ_{NLO}/σ_{LO}) $\approx 1.5 \pm 0.5$

Why is this? Many pseudoscientific explanations

Explosion of # of diagrams ($n_{Diagrams} \approx n!$) New initial states contributing at higher orders (E.g., $gq \rightarrow Zq$) Inclusion of low-x (non-DGLAP) enhancements Bad (high) scale choices at Lower Orders, ...

Theirs not to reason why // Theirs but to do and die

Tennyson, The Charge of the Light Brigade

Changing the scale(s)

Why scale variation ~ uncertainty?

Scale dependence of calculated orders must be canceled by contribution from uncalculated ones (+ non-pert)

$$\alpha_s(Q^2) = \alpha_s(m_Z^2) \frac{1}{1 + b_0 \ \alpha_s(m_Z) \ln \frac{Q^2}{m_Z^2} + \mathcal{O}(\alpha_s^2)}$$

$$b_0 = \frac{11N_C - 2n_f}{12\pi}$$

$$\rightarrow \quad \left(\alpha_s(Q'^2) - \alpha_s(Q^2)\right) |M|^2 = \alpha_s^2(Q^2) |M|^2 + \dots$$

→ Generates terms of higher order, but proportional to what you already have $(|M|^2)$ → a first naive^{*} way to estimate uncertainty

*warning: some theorists believe it is the only way ... but be agnostic! There are other things than scale dependence ...



Complicated final states

Intrinsically <u>Multi-Scale</u> problems with Many powers of α_s

Complicated final states

Intrinsically <u>Multi-Scale</u> problems with Many powers of α_s

 $\begin{array}{|c|c|c|c|c|} \hline \textbf{E.g., W + 3 jets in pp} \\ \hline \alpha_s^3(m_W^2) < \alpha_s^3 \left(m_W^2 + \left\langle p_{\perp}^2 \right\rangle \right) < \alpha_s^3 \left(m_W^2 + \sum_i p_{\perp i}^2 \right) \end{array} \end{array}$

Complicated final states

Intrinsically <u>Multi-Scale</u> problems with Many powers of α_s

E.g., W + 3 jets in pp $\alpha_s^3(m_W^2) < \alpha_s^3 \left(m_W^2 + \langle p_\perp^2 \rangle\right) < \alpha_s^3 \left(m_W^2 + \sum_i p_{\perp i}^2\right)$ <u>Global Scaling</u>: jets don't care about mw $\alpha_s^3(\min[p_\perp^2]) < \alpha_s^3(\langle p_\perp^2 \rangle) < \alpha_s^3(\max[p_\perp^2])$

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p_{⊥1}= 50 GeV p_{⊥2}= 50 GeV p_{⊥3}= 50 GeV

10

9

8

6

5

4

3

2

1

Choice

 α_s Cubed

8

7

6

5

4

3

2

1

 $\frac{10}{\alpha_{s}^{3}}$

10 -3



with Many powers of α_s

E.g., W + 3 jets in pp $a_{s}^{3}(m_{W}^{2}) < \alpha_{s}^{3}(m_{W}^{2} + \langle p_{\perp}^{2} \rangle) < \alpha_{s}^{3}\left(m_{W}^{2} + \sum_{i} p_{\perp i}^{2}\right)$ Global Scaling: jets don't care about mw $a_{s}^{3}(\min[p_{\perp}^{2}]) < \alpha_{s}^{3}(\langle p_{\perp}^{2} \rangle) < \alpha_{s}^{3}(\max[p_{\perp}^{2}])$ MC parton showers: "Local scaling" $\alpha_{s}(p_{\perp 1})\alpha_{s}(p_{\perp 2})\alpha_{s}(p_{\perp 3}) \sim \alpha_{s}^{3}\left(\langle p_{\perp}^{2} \rangle_{geom}\right)$

 $p_{\perp 1}$ = 500 GeV $p_{\perp 2}$ = 100 GeV $p_{\perp 3}$ = 30 GeV

Complicated final states

Intrinsically <u>Multi-Scale</u> problems with Many powers of α_s

If you have multiple QCD scales \rightarrow variation of μ_R by factor 2 in each direction not good enough! (nor is \times 3, nor \times 4) Need to vary also functional dependence on each scale!

